

13.5

a.)

population inversion is required since the cross section for stimulated emission, and absorption is the same.

if $N_1 > N_2$, photons will ~~be not absorbed~~ cause $N_1 \rightarrow N_2$ transitions more often than stimulating emission.

b.)

usually lower is better since in a 4-level system, thermal photons may cause population of N_1 .

13.12

C^{+5} is like Hydrogen atom with $Z=6$.

$$E_n = -13.6 \text{ eV} \frac{Z^2}{n^2} = 490 \text{ eV} \cdot \frac{1}{n^2}$$

$$E = \frac{hc}{\lambda} = \frac{1240 \text{ eV} \cdot \text{nm}}{18 \text{ nm}} = 68.9 \text{ eV}$$

$$n_2 \rightarrow n_1$$

$$n_2 = 3$$

$$n_1 = 2$$

gives correct energy.

13.17

$$\lambda = 633$$

$$L = 0.1 \text{ m}$$

a.) if L changes by $\frac{\lambda}{2}$, this changes the mode by 1

$$\Delta L = \frac{\lambda}{2} = L(8 \times 10^{-6} \text{ K}^{-1}) \Delta T$$

$$\Delta T = \frac{\lambda}{2L(8 \times 10^{-6} \text{ K}^{-1})} = 0.4 \text{ K}$$

b.) as per example 13.2

$$\frac{\Delta \lambda}{\lambda} \approx \sqrt{\frac{2kT}{mc^2}} \quad \text{for neon at } T = 300 \text{ K}$$

$$\Delta \lambda = \lambda \sqrt{\frac{(2)(10258 \text{ eV})}{20 \cdot 1 \times 10^9 \text{ eV}}} = 10^{-3} \text{ nm}$$

number of modes is \sim

$$\Delta q = \frac{2L}{\lambda - \Delta \lambda} - \frac{2L}{\lambda + \Delta \lambda} \approx 1 \text{ mode}$$

Intensity will make small changes as it moves from one mode to another

if $\Delta q \ll 1$ there would be sharp changes in Intensity

and $\Delta q \gg 1$ no changes in Intensity

13.28

$$\lambda = 1.06 \text{ } \mu\text{m}$$

metastable state $\tau_0 = 0.1 \text{ ms}$

$$E_2 - E_1 = \frac{1}{4} \text{ eV}$$

a.) If E_2 is populated by thermal energy, this makes it harder for population inversion.

At 400 K, allowing a 100 C increase over room temp,

$$kT = 0.034 \text{ eV}$$

$$\Delta E = 0.25 \text{ eV}$$

Maxwell Boltzmann factor

$$e^{-\Delta E/kT} = 10^{-3}$$

Cooling not that critical.

b.) The E_3 metastable lifetime does not affect maximum average continuous power. $E_3 \rightarrow E_2$ transitions are by stimulated emission, not spontaneous emission. The life time is for spontaneous emission.

Limiting effect is how fast we can re-populate E_3 after a stimulated transition.

Assume other decays have $\tau \sim 10^{-8} \text{ s}$.

$$P = \frac{E}{t} = \left(\frac{hc}{\lambda} \right) (N) \frac{1}{\tau} = \left(\frac{1240 \text{ eV nm}}{1064 \text{ nm}} \right) \left(\frac{1.6 \times 10^{-19} \text{ J}}{1 \text{ eV}} \right) (10^{17}) 10^8 \text{ s}^{-1}$$

= 2 MW - not realistic, although

KW Nd:YAG is common

12.5

$$E_F = 9.39 \text{ eV}$$

$$n = \frac{8\sqrt{2} \pi m^{3/2}}{h^3} \left(\frac{2}{3} E_F \right)^{3/2}$$

$$n = \frac{8\sqrt{2} \pi (mc^2)^{3/2}}{h^3 c^3} \left(\frac{2}{3} E_F \right)^{3/2}$$

$$= \frac{8\sqrt{2} \pi (0.511 \times 10^6 \text{ eV})^{3/2}}{(1240 \text{ nm}\cdot\text{eV})^3} \left(\frac{2}{3} \cdot 9.39 \text{ eV} \right)^{3/2}$$

$$= 107 \frac{\text{electrons}}{\text{nm}^3}$$

$$\rho = (7.13 \times 10^3 \text{ kg/m}^3) \left(\frac{\text{m}}{10^9 \text{ nm}} \right)^3 \left(\frac{1000 \text{ g}}{1 \text{ kg}} \right) \left(\frac{N_A}{65.4 \text{ g}} \right) = 66 \text{ atoms/nm}^3$$

\uparrow
 65.4 g/mole

gives ~ 1.6 electron/atom

#2.22

$E_g > \frac{hc}{\lambda}$ for visible wavelengths

in Diamond, but not in silicon

Diamond $E_g = 5.5 \text{ eV}$

Si $E_g = 1.1 \text{ eV}$

lowest energy visible ~

$$\frac{1240 \text{ eV}\cdot\text{nm}}{700 \text{ nm}} = 1.7 \text{ eV}$$

highest energy visible

$$\frac{1240 \text{ eV}}{400 \text{ nm}} = 3.1 \text{ eV}$$

12.24

$$V = \pm 1 \text{ V}$$

$$T = 300 \text{ K}$$

Current in direction of $n \rightarrow p$ region

$$I_1 = (I_n + I_p) e^{eV/kT}$$

$p \rightarrow n$

$$I_2 = -(I_n + I_p)$$

$$\frac{I_1}{I_2} = -e^{\pm \frac{eV}{kT}} = 10^{17} \text{ or } 10^{-17}$$